

The background of the slide is a vibrant space scene. On the left, a large portion of the Earth is visible, showing its brown and white surface. The rest of the background is a deep blue space filled with numerous white stars of varying sizes and brightness. In the lower center, there is a smaller, blue-tinted sphere, possibly a planet or moon. The overall lighting is bright and ethereal, with a strong blue hue.

# **Atmospheric Waves**

## **Chapter 6**

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# Part 3: Internal Gravity Waves



# ***Internal Gravity Waves***

(Internal) gravity waves are generated by flow over topography, thunderstorms, other mesoscale systems

(Internal) gravity wave effects are: cloud streets, lenticularis clouds, rapid local temperature and pressure changes, clear air turbulence

They carry energy and momentum into the middle atmosphere (stratosphere, mesosphere)

# *Internal Gravity Waves*

Atmospheric gravity waves can only exist when the atmosphere is stably stratified ( $N^2 > 0$ )

**Buoyancy force is the restoring force for gravity waves**

In an incompressible fluid (e.g. ocean) gravity waves travel primarily in the horizontal plane, since reflection at lower and upper boundary occurs for vertically traveling waves.

In the atmosphere: Due to compressibility and stratification, gravity waves may propagate vertically and horizontally.

# *Topographic Gravity Waves*



**Figure:** Air moving over a mountain on an island triggers waves on the downwind (lee) side of the mountain, indicated by the cloud pattern

# *Examples*

## *Undular Bore*



# Examples

## *Mountain-generated waves (lee waves)*



# Internal Gravity Waves

Consider pure internal gravity waves:  
No rotation, no friction

Can be studied with the basic equations for 2D (x-z plane) motions with Boussinesq approximation (almost incompressible)

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + w \frac{\partial u}{\partial z} + \frac{1}{\rho} \frac{\partial p}{\partial x} = 0$$

$$\frac{\partial w}{\partial t} + u \frac{\partial w}{\partial x} + w \frac{\partial w}{\partial z} + \frac{1}{\rho} \frac{\partial p}{\partial z} + g = 0$$

$$\frac{\partial u}{\partial x} + \frac{\partial w}{\partial z} = 0$$

$$\frac{\partial \theta}{\partial t} + u \frac{\partial \theta}{\partial x} + w \frac{\partial \theta}{\partial z} = 0$$

Four equations in five unknowns



# Internal Gravity Waves

Equations must be closed via an equation of state, such as the potential temperature equation:

$$\theta = T \left( \frac{p_0}{p} \right)^{\frac{R_d}{c_p}} = \frac{p}{\rho R_d} \left( \frac{p_0}{p} \right)^{\frac{R_d}{c_p}} = R_d p^{\frac{c_v}{c_p}} \rho^{-1} p_0^{\frac{R_d}{c_p}}$$

Take logarithms of both sides:

$$\ln \theta = \frac{c_v}{c_p} \ln p - \ln \rho + \text{constant}$$

# IGWs: Derivation (1)

Linearize the basic equations by letting

$$\rho = \rho_0 + \rho'$$

$$p = \bar{p}(z) + p'$$

$$\theta = \bar{\theta}(z) + \theta'$$

$$u = \bar{u} + u'$$

$$w = w'$$

The basic state must satisfy:

$$\frac{d\bar{p}}{dz} = -g\rho_0$$

$$\ln \bar{\theta} = \frac{c_v}{c_p} \ln \bar{p} - \ln \rho_0 + \text{constant}$$

Hydrostatic Balance

Definition of Potential Temperature

# IGWs: Derivation (2)

Linearize the material derivative terms:

$$\begin{aligned}
 \frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + w \frac{\partial u}{\partial z} &= \overset{\text{Constant}}{\cancel{\frac{\partial}{\partial t}(\bar{u} + u')}} + (\bar{u} + u') \overset{\text{Constant}}{\cancel{\frac{\partial}{\partial x}(\bar{u} + u')}} + w' \overset{\text{Constant}}{\cancel{\frac{\partial}{\partial z}(\bar{u} + u')}}), \\
 &= \frac{\partial u'}{\partial t} + \bar{u} \frac{\partial u'}{\partial x} + \cancel{u' \frac{\partial u'}{\partial x}} + \cancel{w' \frac{\partial u'}{\partial z}}, \\
 &= \frac{\partial u'}{\partial t} + \bar{u} \frac{\partial u'}{\partial x} \quad \text{Small} \quad \text{Small}
 \end{aligned}$$

$$\begin{aligned}
 \frac{\partial w}{\partial t} + u \frac{\partial w}{\partial x} + w \frac{\partial w}{\partial z} &= \frac{\partial w'}{\partial t} + \overset{\text{Small}}{\cancel{(\bar{u} + u')}} \frac{\partial w'}{\partial x} + \overset{\text{Small}}{\cancel{w' \frac{\partial w'}{\partial z}}}, \\
 &= \frac{\partial w'}{\partial t} + \bar{u} \frac{\partial w'}{\partial x}
 \end{aligned}$$

# IGWs: Derivation (2)

Linearize the thermodynamic equation:

$$\begin{aligned}
 \frac{\partial \theta}{\partial t} + u \frac{\partial \theta}{\partial x} + w \frac{\partial \theta}{\partial z} &= \frac{\partial}{\partial t} (\overline{\theta} + \theta') + (\bar{u} + u') \frac{\partial}{\partial x} (\overline{\theta} + \theta') + w' \frac{\partial}{\partial z} (\overline{\theta} + \theta'), \\
 &= \frac{\partial \theta'}{\partial t} + \bar{u} \frac{\partial \theta'}{\partial x} + \cancel{u' \frac{\partial \theta'}{\partial x}} + w' \frac{\partial \bar{\theta}}{\partial z} + \cancel{w' \frac{\partial \theta'}{\partial z}}, \\
 &= \frac{\partial \theta'}{\partial t} + \bar{u} \frac{\partial \theta'}{\partial x} + w' \frac{\partial \bar{\theta}}{\partial z} \quad \text{Small}
 \end{aligned}$$

No time dependence
No x dependence

# IGWs: Derivation (3)

Linearize  $\frac{1}{\rho} \frac{\partial p}{\partial z} + g = \frac{1}{\rho_0 + \rho'} \left( \frac{\partial \bar{p}}{\partial z} + \frac{\partial p'}{\partial z} \right) + g$

Remember some calculus:

For  $\frac{\rho'}{\rho_0} \ll 1$   $\frac{1}{\rho_0 + \rho'} = \frac{1}{\rho_0} \frac{1}{\left(1 + \frac{\rho'}{\rho_0}\right)} \approx \frac{1}{\rho_0} \left(1 - \frac{\rho'}{\rho_0}\right)$

Linearization yields

$$\frac{1}{\rho} \frac{\partial p}{\partial z} + g \approx \frac{1}{\rho_0} \frac{\partial \bar{p}}{\partial z} \left(1 - \frac{\rho'}{\rho_0}\right) + \frac{1}{\rho_0} \frac{\partial p'}{\partial z} + g = \frac{1}{\rho_0} \frac{\partial p'}{\partial z} + \frac{\rho'}{\rho_0} g$$


Use  $\frac{1}{\rho_0} \frac{\partial \bar{p}}{\partial z} = -g$

# IGWs: Derivation (3)

Potential temperature equation:

$$\ln \theta = \frac{c_v}{c_p} \ln p - \ln \rho + \text{constant}$$

Linearize via  $\theta = \bar{\theta} + \theta'$ ,  $p = \bar{p} + p'$ ,  $\rho = \rho_0 + \rho'$

  $\ln \left[ \bar{\theta} \left( 1 + \frac{\theta'}{\bar{\theta}} \right) \right] = \frac{1}{\gamma} \ln \left[ \bar{p} \left( 1 + \frac{p'}{\bar{p}} \right) \right] - \ln \left[ \rho_0 \left( 1 + \frac{\rho'}{\rho_0} \right) \right] + \text{constant}$

with  $\gamma = \frac{c_p}{c_v}$

# IGWs: Derivation (3)

Again, remember some calculus:

$$\ln(ab) = \ln a + \ln b$$

$$\ln(1 + \epsilon) \approx \epsilon \quad \epsilon \ll 1$$

Apply these rules to potential temperature equation:

$$\ln \left[ \bar{\theta} \left( 1 + \frac{\theta'}{\bar{\theta}} \right) \right] = \frac{1}{\gamma} \ln \left[ \bar{p} \left( 1 + \frac{p'}{\bar{p}} \right) \right] - \ln \left[ \rho_0 \left( 1 + \frac{\rho'}{\rho_0} \right) \right] + \text{constant}$$

To obtain linearized potential temperature equation:

$$\frac{\theta'}{\bar{\theta}} \approx \frac{1}{\gamma} \frac{p'}{\bar{p}} - \frac{\rho'}{\rho_0} \quad \Rightarrow \quad \rho' \approx -\rho_0 \frac{\theta'}{\bar{\theta}} + \frac{p'}{c_s^2}$$

**Speed of sound**

$$c_s^2 = \frac{\bar{p}\gamma}{\rho_0}$$

# ***IGWs: Derivation (3)***

In gravity waves, density fluctuations due to pressure changes are small compared to those due to temperature changes:

$$\left| \rho_0 \frac{\theta'}{\bar{\theta}} \right| \gg \left| \frac{p'}{c_s^2} \right|$$

Therefore, to a first approximation, the potential temperature equation becomes:

$$\frac{\theta'}{\bar{\theta}} \approx -\frac{\rho'}{\rho_0}$$



# IGWs: Derivation (4)

The linearized equation set for internal gravity waves becomes

$$\left( \frac{\partial}{\partial t} + \bar{u} \frac{\partial}{\partial x} \right) u' + \frac{1}{\rho_0} \frac{\partial p'}{\partial x} = 0 \quad (1)$$

$$\left( \frac{\partial}{\partial t} + \bar{u} \frac{\partial}{\partial x} \right) w' + \frac{1}{\rho_0} \frac{\partial p'}{\partial z} - \frac{\theta'}{\bar{\theta}} g = 0 \quad (2)$$

$$\frac{\partial u'}{\partial x} + \frac{\partial w'}{\partial z} = 0 \quad (3)$$

$$\left( \frac{\partial}{\partial t} + \bar{u} \frac{\partial}{\partial x} \right) \theta' + w' \frac{\partial \bar{\theta}}{\partial z} = 0 \quad (4)$$

# IGWs: Derivation (5)

Eliminate  $p'$  by computing  $\frac{\partial(2)}{\partial x} - \frac{\partial(1)}{\partial z}$

Leads to

$$\left( \frac{\partial}{\partial t} + \bar{u} \frac{\partial}{\partial x} \right) \left( \frac{\partial w'}{\partial x} - \frac{\partial u'}{\partial z} \right) - \frac{g}{\bar{\theta}} \frac{\partial \theta'}{\partial x} = 0 \quad (5)$$

Now eliminate  $u'$  from (5) by computing

$$\frac{\partial(5)}{\partial x} + \left( \frac{\partial}{\partial t} + \bar{u} \frac{\partial}{\partial x} \right) \frac{\partial(3)}{\partial z}$$

Leads to the equation:

$$\left( \frac{\partial}{\partial t} + \bar{u} \frac{\partial}{\partial x} \right) \left( \frac{\partial^2 w'}{\partial x^2} + \frac{\partial^2 w'}{\partial z^2} \right) - \frac{g}{\bar{\theta}} \frac{\partial^2 \theta'}{\partial x^2} = 0 \quad (6)$$

# IGWs: Derivation (5)

Eliminate  $\theta'$  from (6) by computing


$$\left( \frac{\partial}{\partial t} + \bar{u} \frac{\partial}{\partial x} \right) (6) + \frac{g}{\bar{\theta}} \frac{\partial^2 (4)}{\partial x^2}$$

Yields single equation for  $w'$

$$\left( \frac{\partial}{\partial t} + \bar{u} \frac{\partial}{\partial x} \right)^2 \left( \frac{\partial^2 w'}{\partial x^2} + \frac{\partial^2 w'}{\partial z^2} \right) + \frac{g}{\bar{\theta}} \frac{\partial \bar{\theta}}{\partial z} \frac{\partial^2 w'}{\partial x^2} = 0$$

Square of Brunt-Väisälä frequency:  
 $\mathcal{N}^2$  assume to be constant

# IGWs: Derivation (6)


$$\left( \frac{\partial}{\partial t} + \bar{u} \frac{\partial}{\partial x} \right)^2 \left( \frac{\partial^2 w'}{\partial x^2} + \frac{\partial^2 w'}{\partial z^2} \right) + \mathcal{N}^2 \frac{\partial^2 w'}{\partial x^2} = 0$$

Consider **wave-like solutions**:

$$\hat{w} = w_r + iw_i \quad m = m_r + im_i \quad \phi = kx + mz - \nu t$$

$$\begin{aligned} w' &= \text{Re} [\hat{w} \exp(i\phi)] \\ &= \text{Re} [(w_r + iw_i) \exp(i(kx + m_r z - \nu t)) \exp(-m_i z)] \\ &= [w_r \cos(\text{Re}(\phi)) - w_i \sin(\text{Re}(\phi))] \exp(-m_i z) \end{aligned}$$

Horizontal wavenumber  $k$  is real, solution is always sinusoidal

# IGWs: Wavenumbers

$$m = m_r + im_i$$

$$w' = [w_r \cos(\text{Re}(\phi)) - w_i \sin(\text{Re}(\phi))] \exp(-m_i z)$$

The vertical wavenumber  $m$  is complex

- **Real** part  $m_r$  describes the **sinusoidal** variation in  $z$
- **Imaginary** part  $m_i$  describes the **exponential decay or growth** of the wave amplitude in  $z$ , depending on the sign of  $m_i$  (positive / negative).

**If  $m$  is real**, total wave number may be regarded as a vector directed perpendicular to line of constant phase, and in the direction of phase increase.

$$\boldsymbol{\kappa} = k\mathbf{i} + m\mathbf{k} \quad \text{with} \quad k = \frac{2\pi}{L_x} \quad m = \frac{2\pi}{L_z}$$

# IGWs: Dispersion

Insert  $w' = \text{Re} [\hat{w} \exp(i\phi)]$

$$\text{into } \left( \frac{\partial}{\partial t} + \bar{u} \frac{\partial}{\partial x} \right)^2 \left( \frac{\partial^2 w'}{\partial x^2} + \frac{\partial^2 w'}{\partial z^2} \right) + \mathcal{N}^2 \frac{\partial^2 w'}{\partial x^2} = 0$$

**IGW Dispersion Relation**

$$(\nu - \bar{u}k)^2 (k^2 + m^2) - \mathcal{N}^2 k^2 = 0$$

**IGW Frequency**

$$\nu = \bar{u}k \pm \frac{\mathcal{N}k}{\sqrt{k^2 + m^2}} = \bar{u}k \pm \frac{\mathcal{N}k}{\kappa}$$

Relative to the mean wind:

- Positive (+) sign means eastward phase propagation
- Negative (-) sign means westward phase propagation

# IGWs: Frequency

If we let  $k > 0$  and  $m < 0$ , then lines of constant phase  $\phi = (kx + mz)$  tilt eastward with increasing height

The choice of the **positive** root

$$\nu = \bar{u}k \pm \frac{\mathcal{N}k}{\sqrt{k^2 + m^2}} = \bar{u}k \pm \overset{+}{\frac{\mathcal{N}k}{\kappa}}$$

then corresponds to an **eastward and downward** phase propagation (relative to the mean flow).

The zonal and vertical phase speeds (positive root) are

$$c_x = \frac{\nu}{k} = \bar{u} + \frac{\mathcal{N}}{\sqrt{k^2 + m^2}} \quad c_z = \frac{\nu}{m} = \frac{\bar{u}k}{m} + \frac{\mathcal{N}k}{m\sqrt{k^2 + m^2}}$$

# IGWs: Phase Relationships

Solutions for the other perturbation fields  $u'$ ,  $p'$ ,  $\theta'$  can be obtained by substituting  $w'$  back into original linearized equations (let  $k > 0$  and  $m < 0$ )

$$\frac{\partial u'}{\partial x} = -\frac{\partial w'}{\partial z} \implies u' = -\frac{m}{k} w'$$

$$\left(\frac{\partial}{\partial t} + \bar{u} \frac{\partial}{\partial x}\right) u' = -\frac{1}{\rho_0} \frac{\partial p'}{\partial x} \implies p' = \rho_0 \frac{(\nu - k\bar{u})}{k} u'$$

$$\left(\frac{\partial}{\partial t} + \bar{u} \frac{\partial}{\partial x}\right) \theta' = -w' \frac{\partial \bar{\theta}}{\partial z} \implies \theta' = \frac{i}{(k\bar{u} - \nu)} \frac{\partial \bar{\theta}}{\partial z} w'$$

Observe  $u'$ ,  $w'$ ,  $p'$  are exactly in phase, with  $\nu - k\bar{u} > 0$

However  $\theta'$  (or  $T'$ ) out of phase by 1/4 wavelength:

Warm air leads pressure ridge eastward

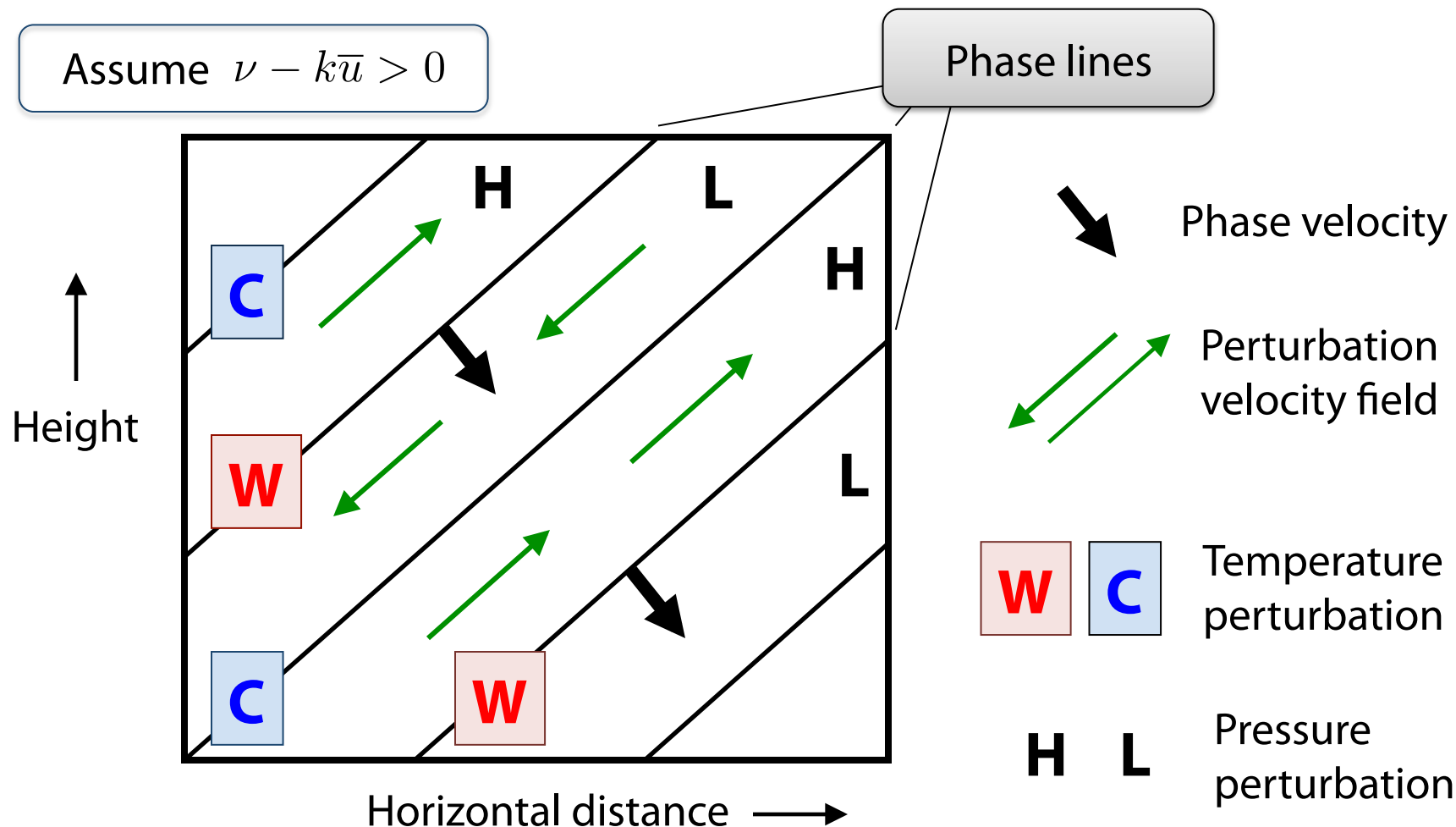
Perpendicular!

Transverse wave:  $\mathbf{k} \cdot \mathbf{u} \equiv k u' + m w' = w' \left[ k \left( -\frac{m}{k} \right) + m \right] = 0$



# IGWs: Cross Section

If we let  $k > 0$  and  $m < 0$ , then lines of constant phase  $\phi = (kx + mz)$  tilt eastward with increasing height



# IGWs: Phase / Group Velocity

The zonal and vertical **phase velocities** are

$$c_x = \frac{\nu}{k} = \bar{u} \pm \frac{\mathcal{N}}{\sqrt{k^2 + m^2}}$$
$$c_z = \frac{\nu}{m} = \frac{\bar{u}k}{m} \pm \frac{\mathcal{N}k}{m\sqrt{k^2 + m^2}}$$

The components of the **group velocity** are

$$c_{g,x} = \frac{\partial \nu}{\partial k} = \bar{u} \pm \frac{\mathcal{N}m^2}{(k^2 + m^2)^{3/2}}$$
$$c_{g,z} = \frac{\partial \nu}{\partial m} = \mp \frac{\mathcal{N}km}{(k^2 + m^2)^{3/2}}$$

Upper and lower signs  
chosen as above.

# IGWs: Phase / Group Velocity

The **vertical** component of **group velocity** has a sign **opposite** to that of the **vertical phase speed**.

That is, **downward phase** propagation implies **upward energy** propagation

$$c_x = \frac{\nu}{k} = \bar{u} \pm \frac{\mathcal{N}}{\sqrt{k^2 + m^2}}$$

$$c_z = \frac{\nu}{m} = \frac{\bar{u}k}{m} \pm \frac{\mathcal{N}k}{m\sqrt{k^2 + m^2}}$$

$$c_{g,x} = \frac{\partial \nu}{\partial k} = \bar{u} \pm \frac{\mathcal{N}m^2}{(k^2 + m^2)^{3/2}}$$

$$c_{g,z} = \frac{\partial \nu}{\partial m} = \mp \frac{\mathcal{N}km}{(k^2 + m^2)^{3/2}}$$

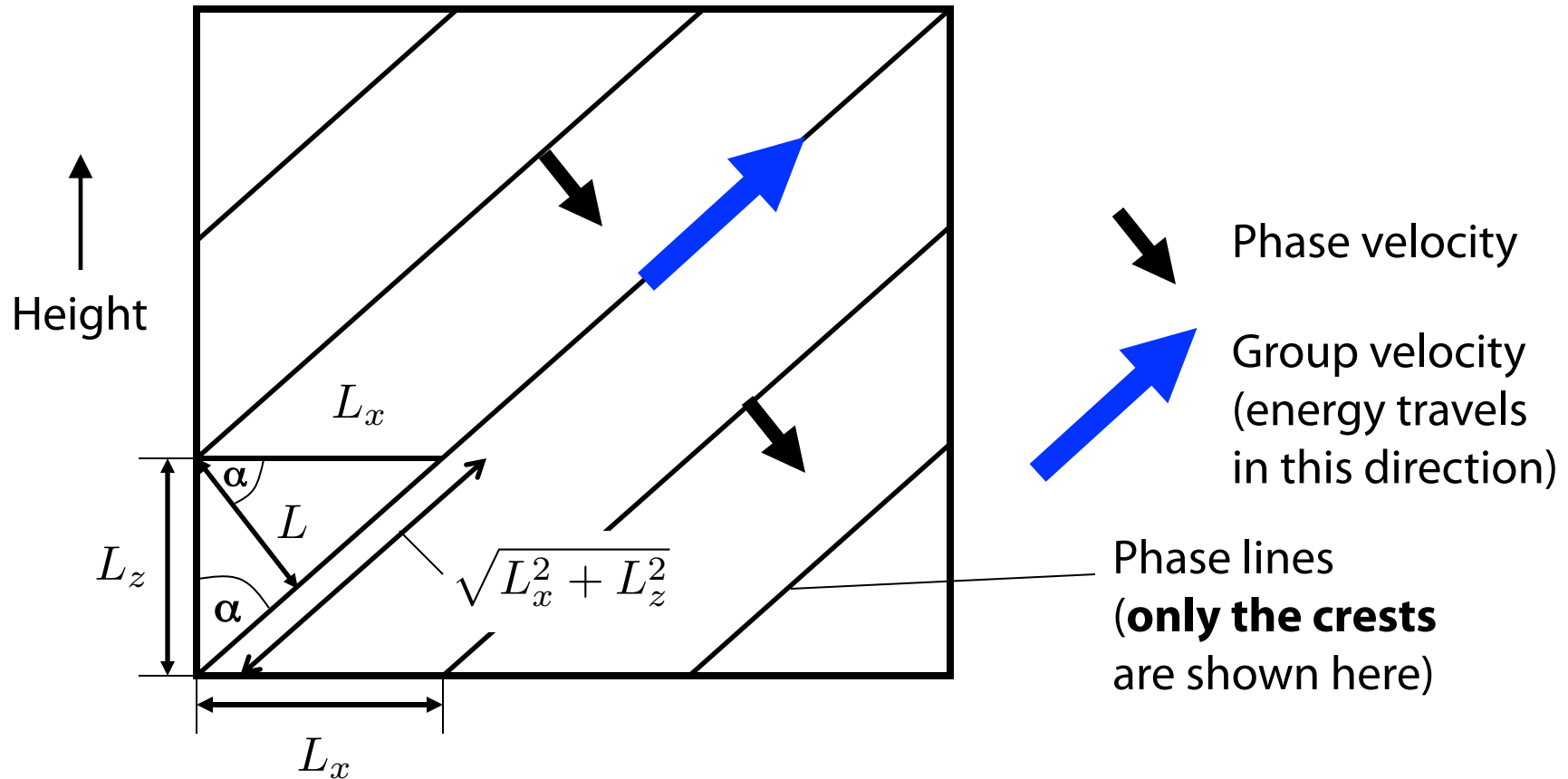
# ***IGWs: Phase / Group Velocity***

Horizontal phase and group velocities point in the same direction (no sign change)

Remarkable property of internal gravity waves:  
Group velocity travels perpendicular to the direction of the phase propagation

That is, in internal gravity waves **energy travels parallel to the crests and troughs** (in contrast to shallow water gravity waves)

# IGWs: Phase / Group Velocity



Angle of phase lines to the local vertical:

$$\cos \alpha = \pm \frac{L}{L_x} = \pm \frac{L_z}{\sqrt{L_x^2 + L_z^2}} = \pm \frac{\left(\frac{2\pi}{m}\right)}{\sqrt{\left(\frac{2\pi}{k}\right)^2 + \left(\frac{2\pi}{m}\right)^2}} = \pm \frac{k}{\sqrt{k^2 + m^2}} = \pm \frac{k}{\kappa}$$

# IGWs: Frequency

Angle of phase lines to the local vertical:

$$\cos \alpha = \pm \frac{L}{L_x} = \pm \frac{L_z}{\sqrt{L_x^2 + L_z^2}} = \pm \frac{\left(\frac{2\pi}{m}\right)}{\sqrt{\left(\frac{2\pi}{k}\right)^2 + \left(\frac{2\pi}{m}\right)^2}} = \pm \frac{k}{\sqrt{k^2 + m^2}} = \pm \frac{k}{\kappa}$$

Frequency of internal gravity waves:

$$\hat{\nu} = \nu - \bar{u}k = \pm \frac{\mathcal{N}k}{\sqrt{k^2 + m^2}} = \pm \frac{\mathcal{N}k}{\kappa}$$

**Intrinsic Frequency**

$$\hat{\nu} \equiv \nu - \bar{u}k = \pm \mathcal{N} \cos \alpha$$

Consequence: gravity wave **intrinsic frequency** (relative to mean flow) must be less than buoyancy frequency

# ***IGWs: Movies***

Propagation of an internal gravity wave packet

[https://www.youtube.com/watch?v=cDsNmnpq9\\_o](https://www.youtube.com/watch?v=cDsNmnpq9_o)

Observed waves in the atmosphere

<http://www.youtube.com/watch?v=RpP62QSJM0g>

# Topographic Gravity Waves

Topographic gravity waves are stationary internal waves driven by topography (a lower boundary condition on  $w$ ).

For stationary waves ( $v=0$ ) this means that the **intrinsic frequency is**  $\hat{\nu} = -\bar{u}k < 0$   
(**westward** relative to the mean flow)

Wavelength  $k$  is determined by the wavelength of the topography (flow follows topography near the ground)

For example, idealized mountain height profile:

$$h(x) = h_M \cos(kx)$$



# Topographic Gravity Waves

Topographic gravity waves are stationary internal waves driven by topography (a lower boundary condition on  $w$ ).

For stationary waves, the dispersion relationship for IGWs

$$\left( \cancel{\frac{\partial}{\partial t}} + \bar{u} \frac{\partial}{\partial x} \right)^2 \left( \frac{\partial^2 w'}{\partial x^2} + \frac{\partial^2 w'}{\partial z^2} \right) + \mathcal{N}^2 \frac{\partial^2 w'}{\partial x^2} = 0$$

simplifies to

$$\left( \frac{\partial^2 w'}{\partial x^2} + \frac{\partial^2 w'}{\partial z^2} \right) + \frac{\mathcal{N}^2}{\bar{u}^2} w' = 0$$

# Topographic Gravity Waves

Seek stationary ( $\nu = 0, \hat{\nu} = \bar{u}k$ ) wave solutions:

$$w' = \text{Re} [\hat{w} \exp(i\phi)] = w_r \cos \phi - w_i \sin \phi$$

$$\hat{w} = w_r + iw_i, \quad m = m_r + im_i, \quad \phi = kx + mz$$

yields the dispersion relationship

$$m^2 = \frac{\mathcal{N}^2}{\bar{u}^2} - k^2$$

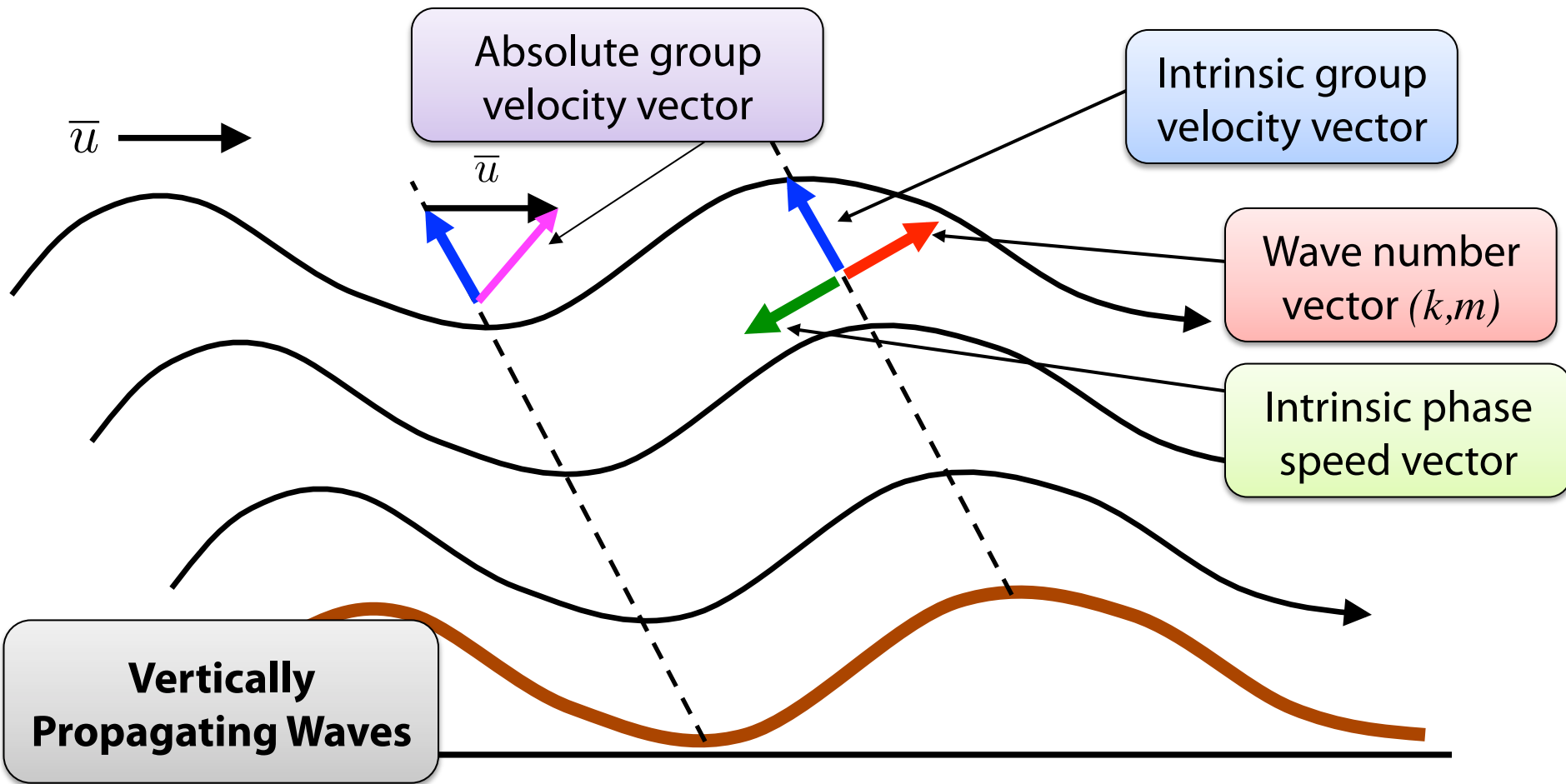
If  $|\bar{u}| < \mathcal{N}/k$ ,  $m^2 > 0$  ( $m$  must be real). Solutions have the form of **vertically propagating waves** (with constant amplitude):

$$w' = \hat{w} \exp(i(kx + mz))$$

Wide (broad) mountain  $\longleftrightarrow$  small  $k$   $\longleftrightarrow$   $m$  real

Ridges and trough lines slope with height, opposite to  $u$

Upward flow of energy and momentum



# Topographic Gravity Waves

If  $|\bar{u}| > \mathcal{N}/k$ ,  $m^2 < 0$  ( $m$  is purely imaginary).

Solutions have the form of **vertically trapped waves** (that decay with height):

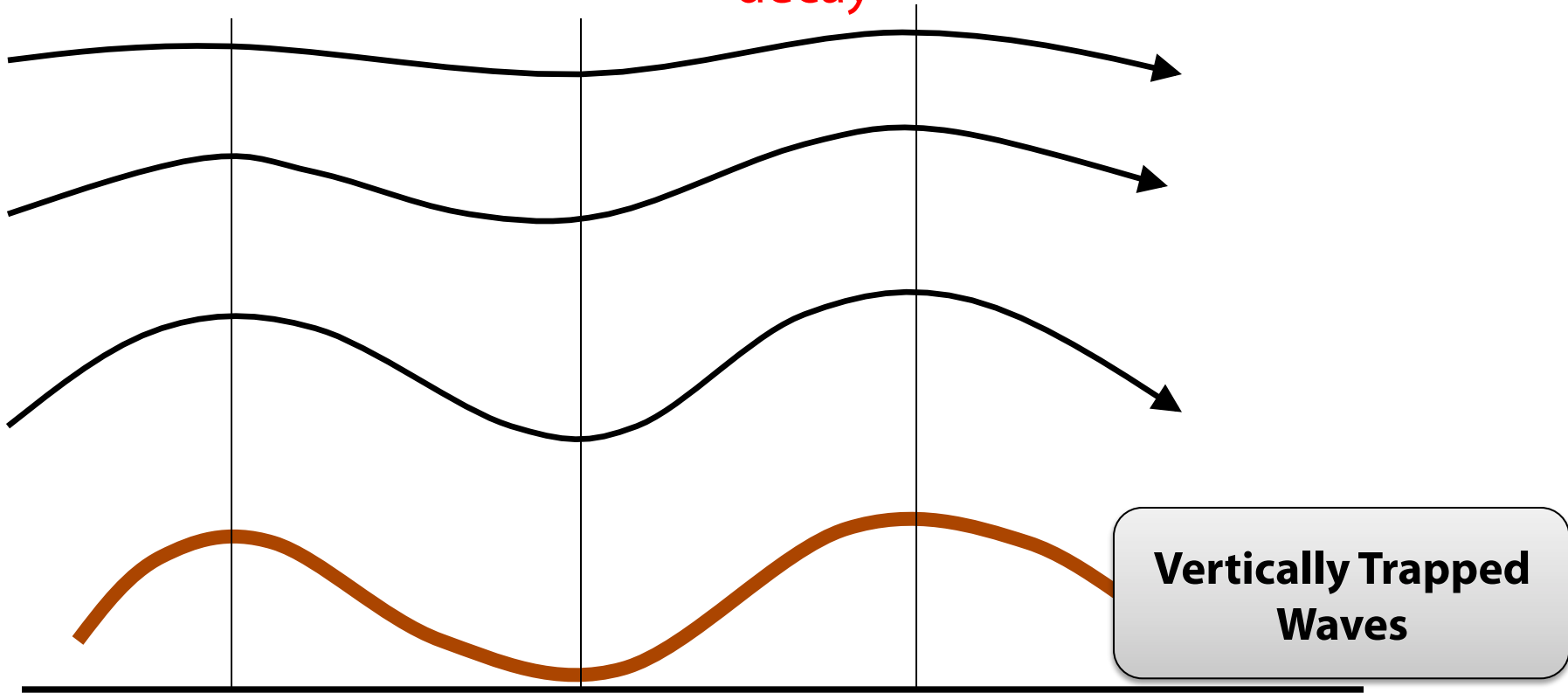
$$m^2 = \frac{\mathcal{N}^2}{\bar{u}^2} - k^2 \quad \rightarrow \quad w' = \hat{w} \exp(ikx) \exp(-m_i z)$$

Narrow mountain  $\longleftrightarrow$  large  $k$   $\longleftrightarrow$   $m$  imaginary

Ridges and trough lines aligned with height

No tilt, no momentum transport

$$w' = \hat{w} \exp(ikx) \underbrace{\exp(-m_i z)}_{\text{decay}}$$



# Topographic Gravity Waves

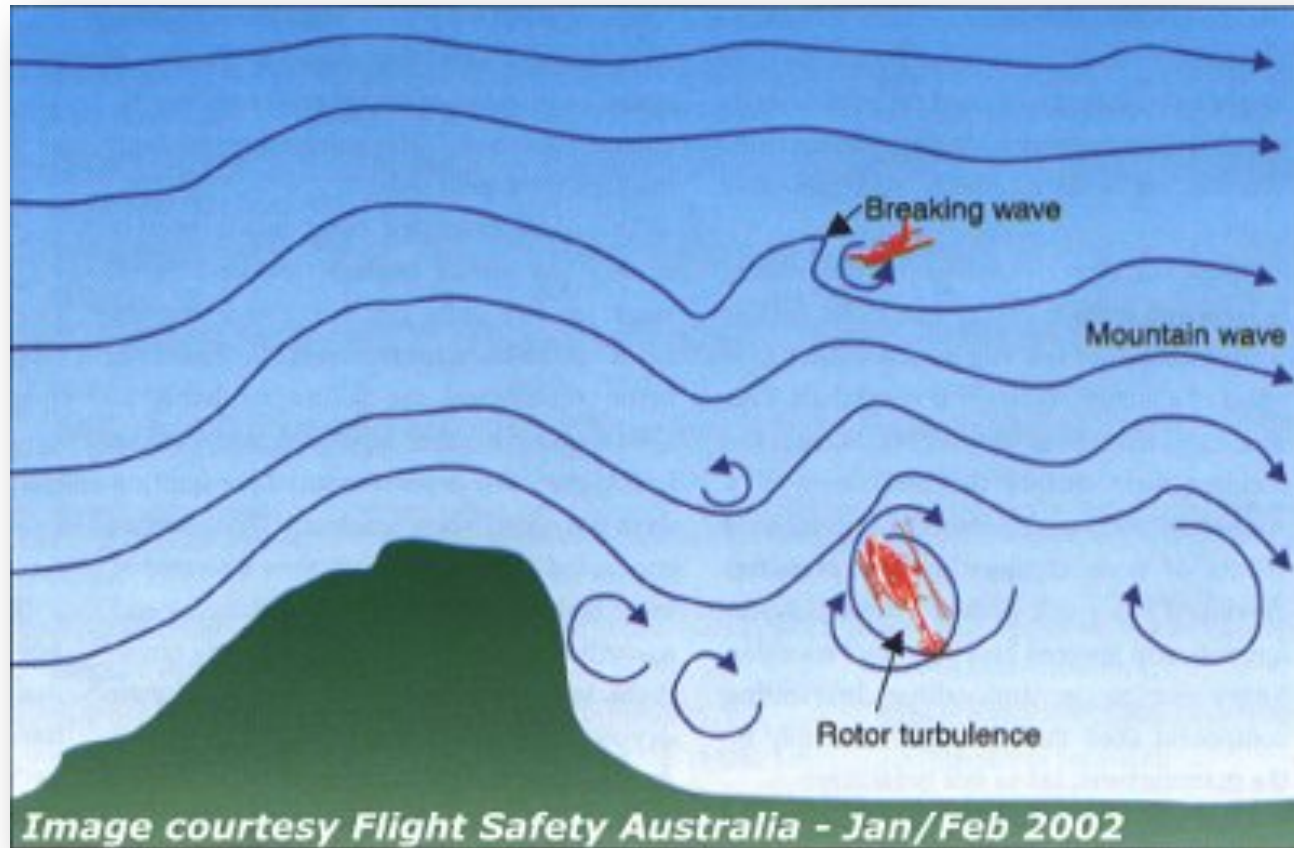
## Summary

**Vertical propagation** of gravity waves is only possible when  $|\bar{u}k|$  is **less than the buoyancy frequency**  $\mathcal{N}$

Stable stratification, wide ridges and relatively weak zonal flow provide favorable conditions for vertically propagating gravity waves.

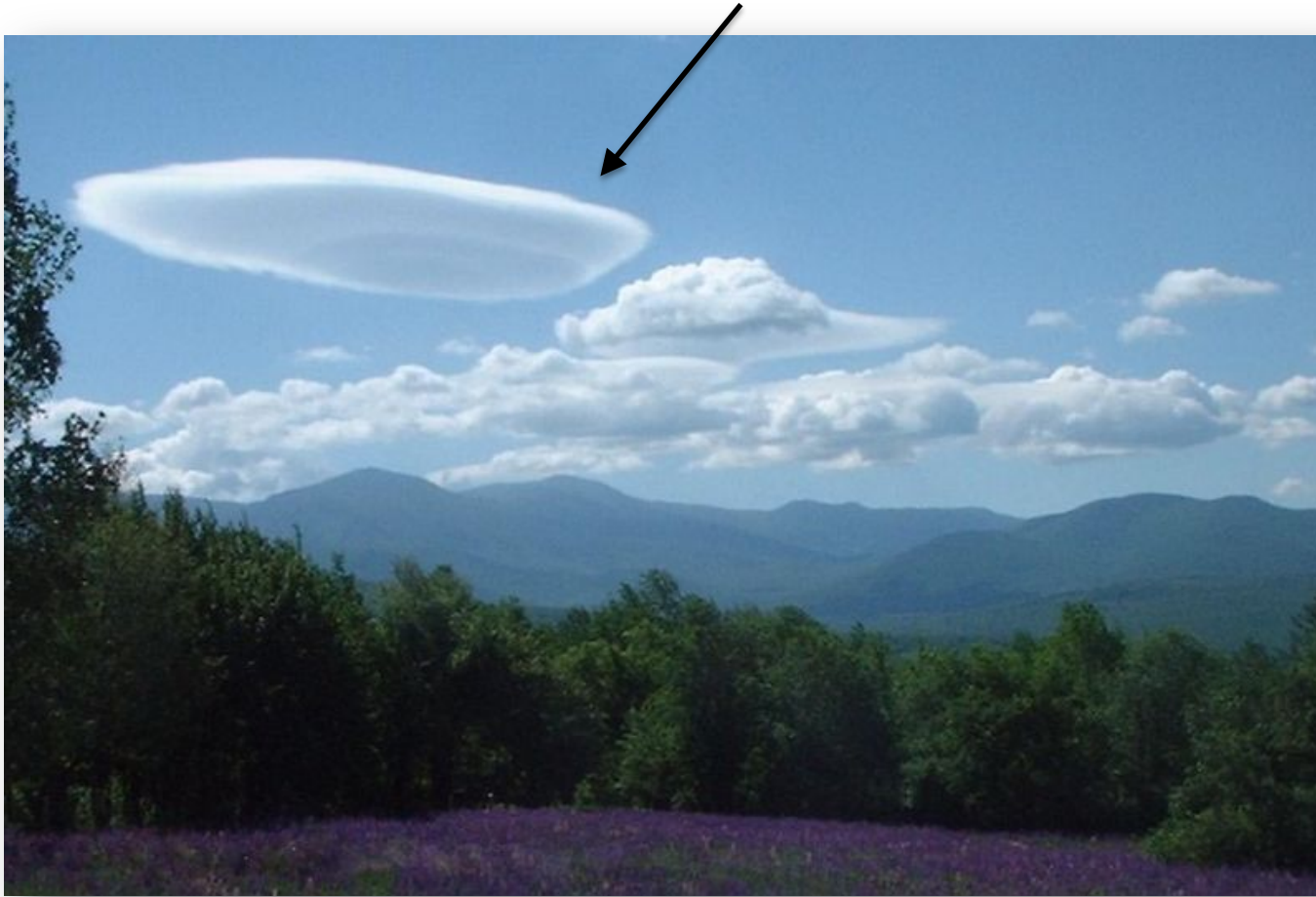
If energy is **transported upward**, phase must be **downward**.

# More Realistic IGWs



**Figure:** Circulation is more complex in real gravity waves:  
For example, rotors, breaking waves, wind shear, varying  $\mathcal{N}$

# *Lenticularis Cloud*



**Figure:**  
Lenticularis cloud  
forming on the  
lee side of the  
mountain

**Movie:** <https://www.youtube.com/watch?v=YKAfKHSeWZc>